## The Quantum and the Continuum : Einstein's Dichotomous Legacies

Talk at : Current Trends in Modern Physics, IISER Kolkata

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## Legacy of the Quantum : Brownian Motion



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## Einstein's formula

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## Einstein's formula

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\left\langle x^{2}\right\rangle_{t}=\frac{k T}{3 \pi \eta r} t
$$

## Implications

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$\square \rightarrow$ Countable degrees of freedom
■ Boltzmann : entropy (irreversibility) $\rightarrow$ probability $\leftrightarrow$ countability
■ Einstein: Explicit proof of existence of discreteness of matter (molecules) and their statistical behaviour

## Legacy of the Quantum : Light Quantum Hypothesis

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Maxwell equations

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\begin{aligned}
\nabla \cdot \vec{E} & =\sum_{l=1}^{N} e_{l} \delta^{(3)}\left(\vec{r}-\vec{r}_{l}(t)\right) \\
\nabla \times \vec{B} & =\sum_{l} e_{l} \vec{v}_{l} \delta^{(3)}\left(\vec{r}-\vec{r}_{l}(t)\right)+\frac{\partial \vec{E}}{\partial t}
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## Einstein :

"... we make use of continuous spatial functions to determine the electromagnetic state of space, so that a finite (countable) number of quantities (charges) cannot be considered as sufficient for the complete determination of the electromagnetic state of space."

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"According to the assumption (hypothesis)..., the energy (of light) ... consists of a finite number of energy quanta localized at points of space that move without dividing, and can be absorbed or generated only as complete units."

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$$
\mathcal{E}=\hbar \omega, \vec{p}=\hbar \vec{k}
$$

## Quantum Electrodynamics : in a flash



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## Quantum Oscillator



## QED (free photons) : Many Uncoupled Oscillators

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\begin{aligned}
a_{\mathrm{e}} & =\left(\frac{g-2}{2}\right)_{\mathrm{e}}^{Q \mathrm{QD}}=0.5 \frac{\alpha}{\pi}-0.32848\left(\frac{\alpha}{\pi}\right)^{2}+1.19\left(\frac{\alpha}{\pi}\right)^{3} \cdots \\
& =(1159652.4 \pm 0.4) \times 10^{-9} \\
a_{\mu} & =\left(\frac{g-2}{2}\right)_{\mu}^{Q D D}=0.5 \frac{\alpha}{\pi}+0.76578\left(\frac{\alpha}{\pi}\right)^{2}+24.45\left(\frac{\alpha}{\pi}\right)^{3} \ldots \\
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But how do we recover Classical $\mathbf{E}$ and $\mathbf{B}$ fields ?

## Quantum to classical oscillator : coherent states

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\begin{aligned}
\hat{a}|\alpha\rangle & =\alpha|\alpha\rangle \\
|\alpha, t\rangle & =\sum_{n=0}^{\infty} \frac{\alpha^{n} e^{-\left[i\left(n+\frac{1}{2}\right) \omega t+\left(|\alpha|^{2} / 2\right)\right]}}{\sqrt{n!}}|n\rangle \\
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Minimum uncertainty states

$$
\Delta x \Delta p=\hbar
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Continuum fields \& classical electrodynamics emerge from QED in the semicl approximation

## Legacy of the Continuum : Galilean Sptm (1+1 dim)



Equivalence class of positions for each instant of time

## Chasing Light : Sp Rel Sptm continuum (1+1 dim)



Hyperboloid: Equivalence class of events

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Hyperboloid: Equivalence class of events
Sptm geometry : non-Euclidean but flat (global)

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$\square \rightarrow$ Spacetime is locally flat but globally curved


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## Gravitational Force originates from Curved Sptm

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Evidence of Dynamical Sptm : Gravitational waves, Expanding Universe, Black Holes

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N SR Point particle energy momentum tensor (classical)

$$
T^{a b}(x)=\sum_{l=1}^{N} m_{l} \int d \tau u_{l}^{a} u_{l}^{b} \delta^{(4)}\left(x-\bar{x}_{l}(\tau)\right)
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GR : Matter $\rightarrow$ Sptm geometry $\rightarrow$ Black Holes GR : Sptm Geometry $\rightarrow$ Matter: Big Bang Both are examples of Sptm singularities in GR!

## Sptm Singularities : Gravitational Collapse

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K<\Delta|\triangle| \ggg>\|+
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Einstein eq. $\Rightarrow$ sptm curvature $\rightarrow \infty$ !!!

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Raychaudhuri Eq : Sptm geometry illdefined at singularity

## Further Conundrum : Black Hole Horizon

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## Area Increase Theorem

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Horizon area can never decrease : Hawking
Analogue of Second Law of thermodynamics: $A_{\text {hor }} \leftrightarrow S$ Mere analogue or more ? If more, microstates ? Black hole : exact solution of Einstein eq. !

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Bekenstein: Microstates necessary for Black Hole Entropy must originate from quantum GR !

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No complete theory yet !
Concrete proposals : Loop Quantum Gravity, Causal Dynamical Triangulations, Spin Foams, ...

## LQG and Black Hole Entropy : Resolution of the Dichotomy

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■ Vertices : invariant $S U(2)$ tensors.
■ Graph : quantum state of space in Spin network basis
■ Geom observables : bounded, discrete spectra

## Area Spectrum

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$$
\begin{aligned}
\hat{\mathcal{A}}_{S} & \equiv \sum_{l=1}^{N} \int_{S_{l}} \operatorname{det}^{1 / 2}\left[{ }^{2} g(\hat{E})\right] \\
a\left(j_{1}, \ldots, j_{N}\right) & =\left.8 \pi \gamma\right|_{P} ^{2} \sum_{p=1}^{N} \sqrt{j_{p}\left(j_{p}+1\right)} \\
\lim _{N \rightarrow \infty} a\left(j_{1}, \ldots . j_{N}\right) & \leq \mathcal{A}_{c l}+O\left(l_{P}^{2}\right) \text { for } j_{p} \leq \frac{k}{2}
\end{aligned}
$$

## Quantum Black Hole (non-rotating)

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## Horizon Description

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■ SU(2) Chern-Simons gauge fields on horizon with punctures carrying spin $j_{I}, I=1, \ldots, N$

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$$
F_{a b}^{i} \Psi=-\frac{k}{2 \pi} \sum_{p} a_{l H, a b}(p) \delta^{(2)}\left(x, x_{p}\right) J_{(p)}^{i} \psi
$$

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Systematic, finite corrections to Bekenstein-Hawking entropy : signature of LQG

## It from Bit

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$A_{\text {plaq }} \sim I_{P I}^{2}: A_{H} / A_{\text {plaq }} \equiv N_{H} \gg 1$

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Virtual ēe pairs disintegrate near the horizon, some drift away Hawking's treatment: Semiclassical!i.e., sptm classical, matter-radiation quantal
If sptm is also quantized, is $\mathbf{b} \mathbf{h}$ radiation still thermal ?

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Recall Horizon Description: CS gauge fields coupled to bulk spinnet (LQG)

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■ What about a formula involving $G, c, \hbar, \Lambda_{Q C D}$ ? Does this occur in Physics ?

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Current Interest : Black hole entropic approach to critical NS mass

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■ Einsteinian sptm continuum is emergent !

